

Cosmología del Big Bang

- * Bases observacionales y teóricas
- * Expansión del Universo
- * Introducción al álgebra tensorial
- * Introducción a la Relatividad General

Bases Observacionales:

- * Expansión del Universo

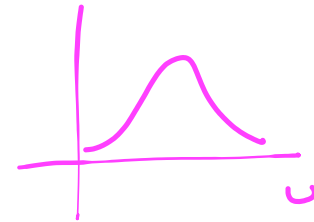
$$v = H d$$

$$H_0 = 74 \text{ km/s Mpc}$$

* La existencia de la radiación cósmica de fondo

$$T = 2.73 \text{ K}$$

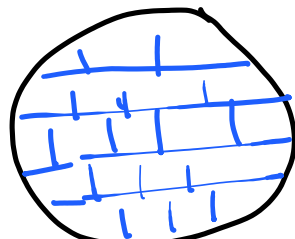
* La nucleosíntesis primordial



Bases teóricas.

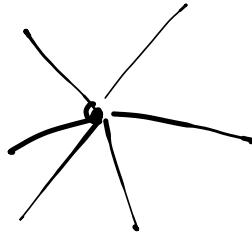
La métrica de FRW:

Homogeneidad

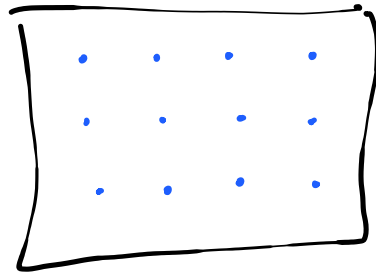


$$10^5$$

Isotropia:



Homogêneo e isotropo



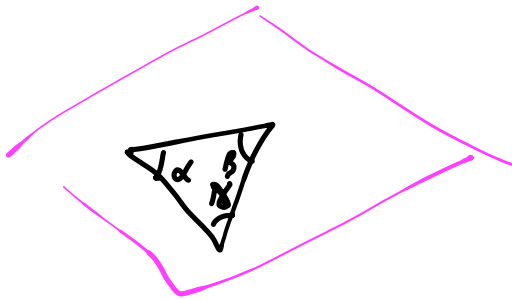
FRW:

$$ds^2 = dt^2 + a^2(t) \left[\frac{dr^2}{1 - kr^2} + r^2 (d\theta^2 + \sin^2\theta d\phi^2) \right]$$

$$0 \leq r < \infty \quad \left[\quad r - r_0 < r^- \right]$$

$$k = \begin{cases} -1 \\ +1 \\ 0 \end{cases} \checkmark$$

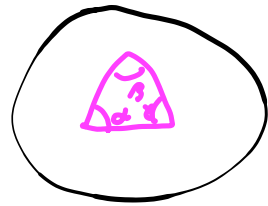
A flat $\therefore k=0$



$$\alpha + \beta + \gamma = \pi$$

$$ds^2 = dx^2 + dy^2$$

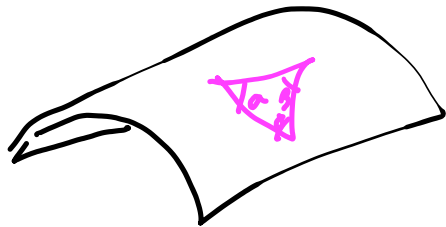
Superficie de una esfera.



$$\alpha + \beta + \gamma > \pi$$

$$ds^2 = dr^2 + R^2 \sin^2(r/R) d\theta^2$$

El hiperboloide:



$$\alpha + \beta + \gamma < \pi$$

$$ds^2 = dr^2 + R^2 \sinh^2(r/R) d\sigma^2$$



* Álgebra Tensorial

* Convención de sumas

$$\sum_n \quad , \quad 1, 2, 3, \dots, n$$

$$\sum_{i=1} a_i = a_1 + a_2 + \dots + a_n$$

$$\sum_{i=1}^n a_i b_i = a_1 b_1 + a_2 b_2 + \dots + a_n b_n$$

$$\sum_{j=1}^n a_{ij} x_j = a_{i1} x_1 + a_{i2} x_2 + \dots + a_{in} x_n$$

* Conversi3n de Einstein.

$$a_i b_i = a_1 b_1 + a_2 b_2 + \dots + a_n b_n$$

$$a_{ij} x_j = a_{i1} x_1 + a_{i2} x_2 + \dots + a_{in} x_n$$

$$a_{ij} x_j = a_{i\ell} x_\ell$$

$i = \text{libre}$

$j = \text{mudo}$

Ejemplo: Si $n=3$, escriba explícitamente la ecuación que está representada por la expresión $y_i = a_{ir} x_r$

$$y_i = a_{i1} x_1 + a_{i2} x_2 + a_{i3} x_3$$

$$y_1 = a_{11} x_1 + a_{12} x_2 + a_{13} x_3$$

$$y_2 = a_{21} x_1 + a_{22} x_2 + a_{23} x_3$$

$$y_3 = a_{31} x_1 + a_{32} x_2 + a_{33} x_3$$

Doble suma:

$$\begin{aligned}
 a_{ij} x_i y_j &= a_{i1} x_i y_1 + a_{i2} x_i y_2 + \dots + a_{in} x_i y_n \\
 &= a_{11} x_1 y_1 + a_{21} x_2 y_1 + \dots + a_{n1} x_n y_1 \\
 &\quad + a_{12} x_1 y_2 + a_{22} x_2 y_2 + \dots + a_{n2} x_n y_2 \\
 &\quad + a_{1n} x_1 y_n + \dots + a_{2n} x_2 y_n + \dots + a_{nn} x_n y_n
 \end{aligned}$$

Ejemplo: si $n=2$, escriba explícitamente las ecuaciones representadas por la expresión $y_i = C_i^r a_{rs} x_s$

$$r = 1, 2$$

$$s = 1, 2$$

$$r: y_i = C_i^1 a_{1s} x_s + C_i^2 a_{2s} x_s$$

$$s: y_i = C_i^1 a_{i1} x_1 + C_i^2 a_{i2} x_2 + C_i^2 a_{21} x_1 + C_i^2 a_{22} x_2$$

$$Y_1 = C_1^1 a_{11} x_1 + C_1^2 a_{12} x_2 + C_1^2 a_{21} x_1 + C_1^2 a_{22} x_2$$

$$Y_2 = C_2^1 a_{11} x_1 + C_2^2 a_{12} x_2 + C_2^2 a_{21} x_1 + C_2^2 a_{22} x_2$$

Delta de Kronecker:

$$\delta_{ij} \equiv \delta_{ji} = \delta^{ij} = \begin{cases} 1, & i=j \\ 0, & i \neq j \end{cases}$$

$$\delta_{ij} = \delta_{ji} \text{ para todo } i, j$$

Ejemplo: Si $n=3$

$$\delta_{ij} x_i x_j = \delta_{11} x_1 x_1 + \delta_{22} x_2 x_2 + \delta_{33} x_3 x_3$$

$$= (x_1)^2 + (x_2)^2 + (x_3)^2$$

$$= x_i x_i$$

$$\delta_{ij} x_i x_j = x_i x_i$$

$$\delta_j^r a_{ir} x_i = a_{ij} x_i$$

* Introducción a la Relatividad General:

La métrica

Es un tensor simétrico covariante de rango 2

$g_{\mu\nu}(x)$ del tipo una métrica

.....

$$ds^2 = g_{\mu\nu}(x) dx^\mu dx^\nu$$

elemento de línea

$$* \quad g_{\mu\nu} = g_{\nu\mu}$$

$$* \quad g = \det(g_{\mu\nu})$$

$$* \quad g_{\mu\nu} g^{\nu\sigma} = \delta_\mu^\sigma$$

$$* \quad T_\mu = g_{\mu\nu} T^\nu$$

$$T^\mu = g^{\mu\nu} T_\nu$$

Theorem: si la matrix Jacobiana de la transformaci3n de un sistema de coordenadas dado (x^i) a un sistema rectangular (x'^i) es $J = \frac{\partial x'^i}{\partial x^j}$, entonces la matrix

recuerda
 $g \equiv (g_{ij})$ del tensor métrico Euclideo en el sistema (x^i)
está dado por:

$$g = J^T J$$

Ejemplos:

* El tensor métrico: la fórmula para la métrica de Minkowski está dada por

$$\eta_{\alpha\beta} = \begin{pmatrix} -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

→ Relatividad especial

Signature: - + + + = +2

$$ds^2 = g_{\mu\nu} dx^\mu dx^\nu$$

$$= g_{00} dx^0 dx^0 + g_{10} dx^1 dx^0 + g_{20} dx^2 dx^0 + g_{30} dx^3 dx^0$$

$$= g_{00} dx^0 dx^0 + g_{11} dx^1 dx^1 + g_{22} dx^2 dx^2 + g_{33} dx^3 dx^3$$

$$= - (dx^0)^2 + (dx^1)^2 + (dx^2)^2 + (dx^3)^2$$

$$x^0 = t$$

$$x^1 = x$$

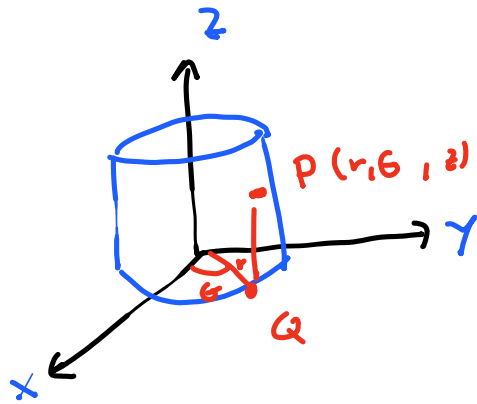
$$x^2 = y$$

$$x^3 = z$$

$$ds^2 = - dt^2 + dx^2 + dy^2 + dz^2$$

Como medir distancias!

* Calcular la métrica en coordenadas cilíndricas:



$$x = r \cos \theta$$

$$y = r \sin \theta$$

$$z = z$$

$$J = \begin{pmatrix} \cos \theta & -r \sin \theta & 0 \\ \sin \theta & r \cos \theta & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$$g = J^T J$$

$$\begin{pmatrix} \cos \theta & \sin \theta & 0 \end{pmatrix} \begin{pmatrix} \cos \theta & -r \sin \theta & 0 \end{pmatrix}$$

$$= \begin{pmatrix} -r \sin \theta & r \cos \theta & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \sin \theta & r \cos \theta & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$$= \begin{pmatrix} 1 & 0 & 0 \\ 0 & r^2 & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$$ds^2 = dr^2 + r^2 d\theta^2 + dz^2$$

* Los símbolos de Christoffel

$$\Gamma_{\mu\nu}^{\alpha} = \frac{1}{2} g^{\alpha\beta} (g_{\beta\mu,\nu} + g_{\beta\nu,\mu} - g_{\mu\nu,\beta})$$

$$\Gamma_{\mu\nu}^{\alpha} = \Gamma_{\nu\mu}^{\alpha}$$

* Tensor de Curvatura: (Riemann tensor)

$$R^{\alpha}_{\beta\mu\nu} = \Gamma^{\alpha}_{\beta\nu, \mu} - \Gamma^{\alpha}_{\beta\mu, \nu} + \Gamma^{\alpha}_{\sigma\mu} \Gamma^{\sigma}_{\beta\nu} - \Gamma^{\alpha}_{\sigma\nu} \Gamma^{\sigma}_{\beta\mu}$$

$$R^{\alpha}_{\beta\mu\nu} = -R^{\alpha}_{\beta\nu\mu}$$

$$R^{\alpha}_{\beta\mu\nu} = 0 \iff$$

El espacio tiempo es plano

* Tensor de Ricci y Escalar de Ricci

o m

Tensor de Ricci: $R_{\alpha\beta} = R_{\alpha\mu\beta}{}^{\mu} = R_{\beta\alpha}$

$$R_{\alpha\beta} = \Gamma_{\alpha\beta,\mu}{}^{\mu} - \Gamma_{\alpha,\mu\beta}{}^{\mu} + \Gamma_{\sigma\mu}{}^{\mu} \Gamma_{\alpha\beta}{}^{\sigma} - \Gamma_{\sigma\beta}{}^{\mu} \Gamma_{\alpha\mu}{}^{\sigma}$$

Escalar de Ricci:

$$R = g^{\mu\nu} R_{\mu\nu}$$

El tensor de Einstein.

$$G^{\mu\nu} = R^{\mu\nu} - \frac{1}{2} g^{\mu\nu} R = G^{\nu\mu}$$

Ejemplo: tensor de Riemann para coordenadas

esféricos

$$R^{\alpha}_{\beta\mu\nu} = \Gamma^{\alpha}_{\beta\nu,\mu} - \Gamma^{\alpha}_{\beta\mu,\nu} + \Gamma^{\alpha}_{\sigma\mu} \Gamma^{\sigma}_{\beta\nu} - \Gamma^{\alpha}_{\sigma\nu} \Gamma^{\sigma}_{\beta\mu}$$

6 componentes que no son idénticamente iguales a cero

$$g_{\mu\nu} = \begin{pmatrix} 1 & & \\ & r^2 & \\ & & r^2 \sin^2 \theta \end{pmatrix}$$

$$g^{\mu\nu} = \begin{pmatrix} 1 & & \\ & 1/r^2 & \\ & & 1/r^2 \sin^2 \theta \end{pmatrix}$$

$$\Gamma_{22}^1, \Gamma_{33}^1, \Gamma_{12}^2 = \Gamma_{21}^1, \Gamma_{33}^2$$

$$\Gamma_{13}^3 = \Gamma_{31}^3, \Gamma_{23}^3 = \Gamma_{32}^3$$

$$\begin{aligned} \Gamma_{22}^1 &= \frac{1}{2} g^{10} (g_{02,2} + g_{12,2} - g_{22,0}) \\ &= \frac{1}{2} g^{11} (g_{12,2} - g_{22,1}) \\ &= -\frac{1}{2} g^{11} g_{22,1} \\ &= -\frac{1}{2} \frac{\partial}{\partial r} (r^2) \end{aligned}$$

$$= -\frac{1}{2} (2r)$$
$$= -r \quad \cancel{f}$$

$$\int_{33}^1 = -r \sin^2 \theta$$

$$\int_{12}^2 = \int_{21}^2 = \frac{1}{r}$$

$$\int_{33}^2 = -\sin \theta \cos \theta$$

$$\int_{13}^3 = \int_{31}^3 = \frac{1}{r}$$

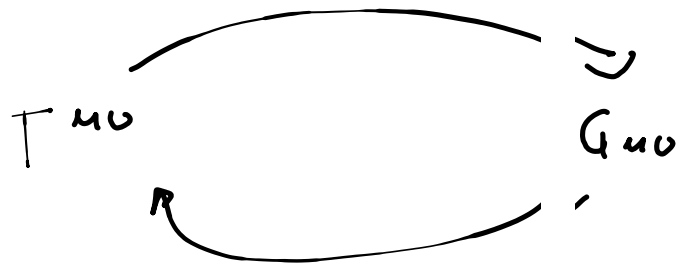
$$\int_{23}^3 = \int_{32}^3 = \cot \theta$$

$$\begin{aligned}
R^1_{212} &= \Gamma^1_{22,1} - \cancel{\Gamma^1_{21,2}} + \frac{\Gamma^1_{\sigma 1} \Gamma^{\sigma}_{22} - \Gamma^1_{\sigma 2} \Gamma^{\sigma}_{21}}{\dots} \\
&= \Gamma^1_{22,1} - \cancel{\Gamma^1_{21,2}} + \Gamma^1_{11} \cancel{\Gamma^1_{22}} + \Gamma^1_{22} \cancel{\Gamma^1_{22}} \\
&\quad + \Gamma^1_{31} \cancel{\Gamma^1_{22}} - \cancel{\Gamma^1_{12}} \Gamma^1_{21} - \Gamma^1_{22} \Gamma^1_{21} - \cancel{\Gamma^1_{32}} \Gamma^1_{21} \\
&= \frac{d}{dr}(-r) + r\left(\frac{1}{r}\right) \\
&= -1 + 1 \\
&= 0
\end{aligned}$$

$$R^{\alpha}_{\beta\mu\nu} = 0 \downarrow$$

* Ecuaciones de campo de Einstein

$$G^{\mu\nu} = 8\pi G T^{\mu\nu}$$



$$G = c = 1$$

$$G^{\mu\nu} = 8\pi T^{\mu\nu}$$

$$1 = \frac{G}{c^2} = 7.425 \times 10^{-27} \text{ m/kg}$$

Constante	S2 u valor	Valor geometrizadas
c	$2.998 \times 10^8 \text{ m/s}$	1
G	$6.674 \times 10^{-11} \text{ m}^3/\text{kg s}^2$	1
m_e	$9.109 \times 10^{-31} \text{ kg}$	$6.764 \times 10^{-58} \text{ m}$
m_p	$1.673 \times 10^{-27} \text{ kg}$	$1.673 \times 10^{-54} \text{ s}$